#### **My Thoughts on MiniBooNE**



## A Review of the MiniBooNE Excess

- 1. Short history of MiniBooNE.
- 2. Excess is EM-like, probably not a background
- A look at key elements of the excess in neutrino, anti-neutrino, and beam dump mode energy, angle, radius, timing, etc.
- 4. Is it neutrino or meson flux related?
- 5. What does the beam off target run tell us.
- 6. Anomaly within the anomaly beam direction hot spot?!
- 7. My thoughts....



#### MiniBooNE (2003-2018) was designed to test the LSND signal



Neutrino mode: search for  $v_{\mu} \rightarrow v_e$  appearance with 18.75E20 POT  $\rightarrow$  assumes CP/CPT conservation Antineutrino mode: search for  $\overline{v_{\mu}} \rightarrow \overline{v_e}$  appearance with 11.27E20 POT  $\rightarrow$  direct test of LSND Beam dump mode: exotic production from proton/neutral-pions: 1.8E20 POT  $\rightarrow$  dark matter search, other.

#### FNAL has done a great job delivering beam over 15 years!

#### 2. Neutrino beam Stable over 15-year run!





3. Events in the Detector

1520 8" PhotoMultipler Tubes (PMT)

2007 data 6.46x10<sup>20</sup> POT

2017 data 6.38x10<sup>20</sup> POT

2019 data 5.9x10<sup>20</sup> POT

160

 $\pi^0$  invariant mass [MeV/c<sup>2</sup>]

180

x° invariant mass [MeV/c<sup>2</sup>

140

#### Magnetic focusing horn



800-ton mineral oil Cerenkov detector 541 m downstream from horn





MB detector is still there, could be turned on again...

#### ~20 Years of Successful MiniBooNE Running and Results!



#### **Robust Excess!**

	ν mode	$\overline{oldsymbol{ u}}$ mode	Combined
Data	2870	478	3348
Unconstr. Background	2322	398.2	2720.0
Constr.	2309	398.7	2707.7
Excess	$560.6 \pm 119.6 \\ 4.8\sigma$	$79.3 \pm 28.6$ $2.8\sigma$	639.9 <u>+</u> 123.0 <b>5.2σ</b>
0.26% (LSND) $\nu_{\mu} \rightarrow \nu_{e}$	676.3	100.0	776.3





R. T. Thornton – Recent Results from MiniBooNE

#### **Neutrino/Antineutrino Excesses**

#### most backgrounds constrained by high stats $v_{\mu}$ rate (poor mans near detector)



Phys Rev D 103 (2021) 5, 052002

#### Excess EM-like $\gamma$ /electron, or co-linear $\gamma\gamma$ /e+e-. No information on hadronic final states



# MiniBooNE Background Radial Distributions: $\pi^0$ misID disfavored

![](_page_8_Figure_1.jpeg)

 $\gamma$ /electron (excess) efficiency decreases near wall due to ~meter extended track length

 $\pi^{0}$  decay lose one gamma-ray at detector edge creating mis-ID electron-like event.

![](_page_8_Figure_4.jpeg)

Excess falls off at wall more rapidly due to lower efficiency. Tighter radial cuts increases excess significance

Selection	Data	Background	Excess	Significance
$200 < E_{\nu}^{QE} < 1250 \ {\rm MeV} \ \& \ R < 5m$	2870	$2309.4\pm119.6$	$560.6 \pm 119.6$	$4.7\sigma$
$150 < E_{\nu}^{QE} < 1250~{\rm MeV}$ & $R < 5m$	3172	$2560.4\pm131.5$	$611.6 \pm 131.5$	$4.7\sigma$
$200 < E_{\nu}^{QE} < 1250~{\rm MeV}$ & $R < 4m$	1978	$1519.4\pm81.9$	$458.6\pm81.9$	$5.6\sigma$
$200 < E_{\nu}^{QE} < 1250~{\rm MeV}$ & $R < 3m$	864	$673.9 \pm 41.2$	$190.1\pm41.2$	$4.6\sigma$

#### MiniBooNE Background Radial Distributions: Dirt and pi0 disfavored

![](_page_9_Figure_1.jpeg)

TABLE IV: The result of log-likelihood shape-only fits to the radial distribution in neutrino mode, assuming only statistical errors, where different processes are normalized to explain the observed event excess. The two-neutrino hypothesis fits the radial distribution best with a  $\chi^2 = 8.4/9$ ndf, while the NC  $\pi^0$  hypothesis has a worse fit with a  $\chi^2 = 17.2/9$ ndf. Also shown is the multiplicative

factor that is required for each hypothesis to explain the observed event excess.

	Hypothesis	Multiplicative factor	$\chi^2/9ndf$
3	NC $\Delta \to N \gamma$ Background	3.18	10.0
1	External Event Background	5.98	44.9
	$\nu_e \ \& \ \bar{\nu}_e$ from $K^0_L$ Decay Background	7.85	14.8
	$\nu_e \ \& \ \bar{\nu}_e$ from $K^\pm$ Decay Background	2.95	16.3
	$\nu_e \ \& \ \bar{\nu}_e$ from $\mu^\pm$ Decay Background	1.88	16.1
	Other $\nu_e~\&~\bar{\nu}_e$ Background	3.21	12.5
	NC $\pi^0$ Background	1.75	17.2
13	Best Fit Oscillations	1.24	8.4

# Recent MicroBooNE Results (arXiv 2110.00409 + 2110.14054 + other uE papers) confirms MiniBooNE backgrounds

New MicroBooNE results demonstrates MB excess is robust – confirmed Delta radiative,  $\pi^0$ , and intrinsic  $v_e$  backgrounds estimates.

![](_page_10_Figure_2.jpeg)

### Low Energy Roll Over: Not run-away backgrounds!

![](_page_11_Figure_1.jpeg)

12

### **Other Background Tests/Studies**

![](_page_12_Figure_1.jpeg)

Different Pi0 generators

Figure 12. Predicted distributions of NC $\pi^0$  background events in MiniBooNE (neutrino mode) as a function of the reconstructed neutrino energy,  $E_{\nu}^{\text{QE}}$ , from different Monte Carlo generators. In each case, we derive the  $\pi^0/e^-$  separation cut using our NUANCE (the generator used by MiniBooNE) MC samples and apply them to the various generators. The three panels correspond to the three different cut shapes introduced in section 3.1. The official MiniBooNE background prediction [2] is shown in red.

arXiv: 2210.08021, Kelly&Kopp

![](_page_12_Figure_5.jpeg)

Table 2. Minimum  $\chi^2$ , number of degrees of freedom (NDF) and Goodness of Fit (GoF) of the analyses of the data of short-baseline  $\overleftarrow{\nu_{\mu}} \rightarrow \overleftarrow{\nu_{e}}$  experiments discussed in the text without (MB) and with ( $\widetilde{\text{MB}}$ ) our enhanced single- $\gamma$  background in MiniBooNE.  $\Delta m_{41}^{2(\text{bf})}$  and  $\sin^2 2\vartheta_{e\mu}^{(\text{bf})}$  are the best-fit values of the corresponding oscillation parameters.

#### MicroBoonE and MiniBooNE 3+1 combined fit: MiniBooNE excess dominates

![](_page_13_Figure_1.jpeg)

Phys. Rev. Lett. 129, 201801 (2022)

![](_page_13_Figure_3.jpeg)

#### ~nsec timing wrt beam: Excess events travel at/near speed of light

High stats muon neutrino data matches 53MHz RF beam structure

![](_page_14_Figure_2.jpeg)

FIG. 14. Illustration of the timing difference between an electron event and a backward-going photon.

#### Very forward pointing angle wrt beam: Elastic Scattering

No funny business from muon neutrino magnetic moment, NC dark matter scattering, etc

![](_page_15_Figure_2.jpeg)

FIG. 16: The MiniBooNE neutrino mode  $\cos \theta$  distribution for  $\cos \theta > 0.9$ , corresponding to the total  $18.75 \times 10^{20}$  POT neutrino data in the visible energy range from 150 to 1250 MeV, for  $\nu_e$  CCQE data (points with statistical errors) and predicted backgrounds (colored histograms). Neutrino-electron elastic scattering events are shown as the hatched region in the "Others" category.

Phys Rev D 103 (2021) 5, 052002

#### **Neutrino Angular (wrt beam) distributions**

Phys Rev D 103 (2021) 5, 052002

![](_page_16_Figure_2.jpeg)

#### Off forward $e/\gamma$ scattering via >MeV Mediator

![](_page_17_Figure_1.jpeg)

N

N

Many other examples will be presented at this work-shop.

#### MiniBooNE and LSND consistent with oscillation or beam dump interpretation

![](_page_18_Figure_1.jpeg)

- MiniBooNE L/E consistent with LSND excess => oscillations?
- MiniBooNE E<sub>excess</sub>/E<sub>beam</sub> same as LSND => beam dump particle production?<sup>19</sup>

#### **Rare Meson Decay Solutions (Meson Portal)**

![](_page_19_Figure_1.jpeg)

Excess rates
between
nu/antinu and
beam-dump
scale with
neutrino flux <b>O</b> R
charged meson
rates.

Phys. Rev. Lett. 129 (2022) 11, 111803 and arXiv:2309.02599

Vector-portal dark matter						
Scenario	$(m_V)$	$m_1, m_{V_2}, m_\chi, m_\chi$	$(\chi')$	$\epsilon_1\epsilon_2 g_2'^2/(4\pi)$	$\chi^2/{ m dof}$	
Single	(17	(, -, 8, 40) Me	V	$3.6 imes10^{-9}$	2.5(2.9)	
Double	(17,	200, 8, 50) M	eV	$1.3  imes 10^{-7}$	2.2(2.6)	
Long-lived (pseudo)scalar						
Scenario $(m_{Z'}, m_{\phi/a})$			$g_{\mu}g$	$g_n \lambda \; [{ m MeV}^{-1}]$	$\chi^2/{ m dof}$	
Scalar		$(49, 1) { m MeV}$	$2.2  imes 10^{-8}$		2.0(2.1)	
Pseudoscalar		(85,1) MeV	Ę	$5.9 \times 10^{-7}$	2.0(2.1)	

TABLE I. Summary of example fits. In the single-mediator scenario,  $m_{V_2}$  is irrelevant, and  $\epsilon_2 = \epsilon_1$  and  $g'_2 \rightarrow g'_1$ . Due to the mass values of the mediators appearing in the scattering process, we fit the data in the limit of nucleon (nucleus) scattering for the double-mediator scenario (the others). The  $\chi^2$  in the parentheses are the values with statistics only.

Channel	limit Model		$(\times 10^{-12})$	Model ii) $(\times 10^{-8})$	
(BR)	$(\times 10^{-8})$	Single	Double	$\phi$	a
$K \to \mu \nu_{\mu} V(\phi)$ [78]	2000 (300)	500	680	230	100
$K \to e \nu_e \nu \nu $ [53]	6000	530	720	_	-
$K \to \mu(e) \nu_{\mu(e)} ee$ [53]	7.4(2.7)	500(530)	680(720)	-	-
$\pi \to \mu(e) \nu_{\mu(e)} X$ [79]	600(50)	0.12(25)	0.17(34)	120(-)	1.1(-)
$\pi \to \mu(e) \nu_{\mu(e)} ee$ [53]	-(0.37)	0.12(25)	0.17(34)	_	-

TABLE II. Relevant exotic decays of  $\pi^{\pm}/K^{\pm}$  and existing upper limits at 90% CL. X stands for invisibly decaying (massive) bosons. The predicted BRs (third though last columns) are based on the following parameter choices:  $(\epsilon_1, \frac{g_1'^2}{4\pi}) \simeq (6.0 \times 10^{-5}, 1)$  for the single-mediator scenario,  $(\epsilon_1, \epsilon_2, \frac{g_2'^2}{4\pi}) \simeq (7.0 \times 10^{-5}, 1.0 \times 10^{-4}, 0.5)$  for the double-mediator scenario,  $(g_{\mu}, g_n, \lambda) \simeq (5 \times 10^{-3}, 10^{-2}, 4.4 \times 10^{-4} \text{ MeV}^{-1})$  for the scalar scenario, and  $(g_{\mu}, g_n, \lambda) \simeq (10^{-2}, 10^{-2}, 6.5 \cdot 10^{-3} \text{ MeV}^{-1})$  for the pseudo-scalar scenario.

#### MiniBooNE Beam Dump: First dedicated proton accelerator Dark Matter search

Phys. Rev. Lett. 118, 221803 (2017), and Phys. Rev. D 98, 112004 (2018)

![](_page_20_Figure_2.jpeg)

Neutrino flux and backgrounds reduced by a factor of x27 Expect ~2 excess events assuming neutrino scaling

![](_page_20_Figure_4.jpeg)

Oscillation cuts: Expect 8 events, observe 6: If excess scaled with POT, or neutral pion production/decay, then expect ~35 events

![](_page_20_Figure_6.jpeg)

#### Anomaly in the Anomaly: Spatial Hot Spot along beam direction Oh no, not more weirdness!!

![](_page_21_Figure_1.jpeg)

Look else-where effect lessens significance, but what about look-here!

![](_page_21_Figure_3.jpeg)

High stats muon/piO samples show spatial anomalies.

#### X-Y plot (perpendicular to the z-axis – beam direction) 11.27E20POT Antineutrino Oscillation-cut Data Set (2006-2012)

![](_page_22_Figure_1.jpeg)

- Efficiency correct background + 2-D Gaussian fit= 55 +/- 22 events,
- Hot spot Gaussian width= ~93 cm (1-sigma) = 2 mRadians
- BPM projection of the Proton Beam direction (the white circle)
   X= 14.2 +/- 35.0 cm

Y= -187.8 +/- 35.0 cm  $\leftarrow$  Detector built 2m higher than beam line

Hot spot spatial test wrt beam: ~3-sigma

![](_page_22_Figure_7.jpeg)

# The real reason for the beam dump mode run was to check the hot spot:

#### From POT scaling expect ~7 oscillation-cut events in hot spot

![](_page_23_Figure_2.jpeg)

- Green dot/dashed line represents SWIC corrected projected beam position and 2σ HotSpot radius X<sub>beam</sub>= 75 cm; Y<sub>beam</sub>= -114cm
- Blue dot and dashed line are antinu best fit hot spot position and 2σ radius: X<sub>Antinu</sub>= 14cm; Y<sub>Antinu</sub>= -188cm

Drell-Yan production of new particle? See Doo-Jin Kim's talk...

### My Thoughts...

- LSND and MiniBooNE appearance signals are robust and have stood the test of time and scrutiny. Solar neutrino anomaly took ~30 years to solve!
- MB excess EM-like, near speed of light propagation (<50 MeV). LSND and MB excesses have similar L/E (oscillations), and similar E<sub>excess</sub>/E<sub>beam</sub> (dump production).
- 3+1 oscillation energy and angle fits not ideal, more complicated oscillations could be at play suggested by global fits or multiple process at work? e.g. sterile neutrinos + meson portals, etc??
- Neutrino SBL experiments provide some of the best rare pion decay measurements (meson portal models)
   is there a signal? uB, CCM, SBND and ICARUS will test in detail soon.
- If MiniBooNE hot spot real (yeah, sure!), then SBND could potentially see a dramatic spatial distribution (~cm hot spot along beam direction), but due to high backgrounds best to run in beam dump mode.
  - Kicker magnet installed in BNB allows running neutrino and beam dump mode simultaneously (nu:BD ratio 5:1, for example).
- Coming years will be interesting and telling, have fun!

#### Source of Unconstrained Errors

energy range.	
Systematic Uncertainty	Fraction of Event Excess
Cross Section	35%
Optical Model	23%
$\pi^+$ Production	14%
Neutrino Flux	7%
$K^0$ Production	4%
$K^+$ Production	4%

TABLE III: The fractional unconstrained systematic uncertainties in the  $200 < E_{\nu}^{QE} < 1250 \text{ MeV}$ 

## **Beam Off Target Running**

![](_page_26_Figure_1.jpeg)

- $\pi^0$  and  $\eta$  produced by protons in the iron quickly decay producing WIMPs ( $\chi$ )
- Charged mesons are absorbed in the iron before decaying, which significantly reduces the neutrino flux (still some production from proton-Air interactions).

#### Kinematics of Events inside and outside the fitted hot spot AntiNeutrino EnuQE: 10.87E20 POT

![](_page_27_Figure_1.jpeg)

200 < E < 475 Data-MC= 68-43.9= 24.1 475 < E < 1250 Data-MC= 68-44.9= 23.9 200 < E < 475 Data-MC= 181-152.5= 28.5 475 < E < 1250 Data-MC= 145-147.6= -2.6

Rcut > 2.0\*Sigma region looks like the neutrino low energy excess in magnitude and shape

## HotSpot Fit Model

#### (Same as for Antinu Data Fits)

- Symmetric 2D Gaussian 4 parameters
  - Gaussian normalization
  - Gaussian width (sometimes fixed)
  - Gaussian X and Y (sometimes fixed).
- Monte Carlo XY background shape from neutrino EM box (no dirt).
  - 1 parameter for normalization
- Data bin errors are +/-sqrt(N<sub>ii</sub>)
- 6x6 binning (good for low stats)
- LogLiklihood fit

	х- Х-ро	sition	. v- Y-positi	on			H2XY	'6bin
	500-		,,				Entries Mean X	130141
Ζ	400						Mean y RMS x	-4.221 199.6
	300							
	200							
	100							
	₀⊨							
	-100							
	-200							
	-300							
	-400	•						
	-500 -500	-400	-300 -200	-100	0 100	200 300	400	 500

#### Alignment Offsets: Working Closely with BNB and Alignment Group

![](_page_29_Figure_1.jpeg)