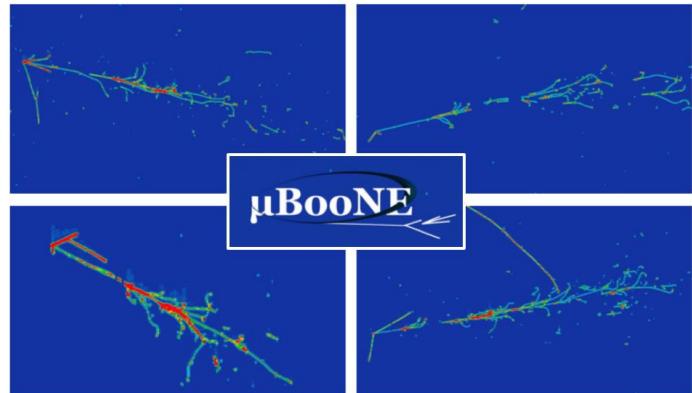


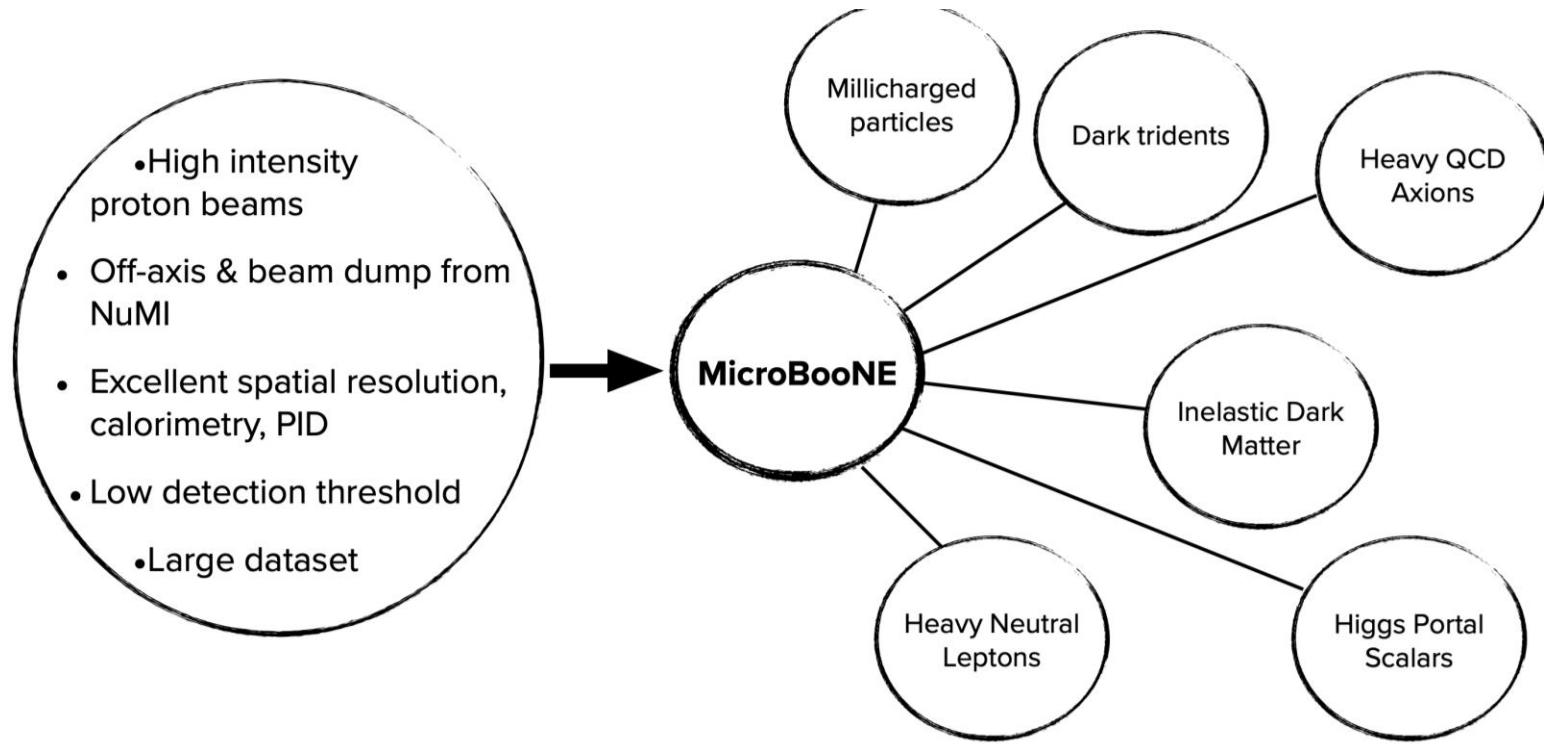
MicroBooNE's BSM physics programme



Justin Evans
University of Manchester



BSM physics with MicroBooNE



BSM physics with MicroBooNE

Results shown in this talk:

arXiv:2312.13945

Phys. Rev. Lett. **132**, 041801 (2024)

arXiv:2308.03924

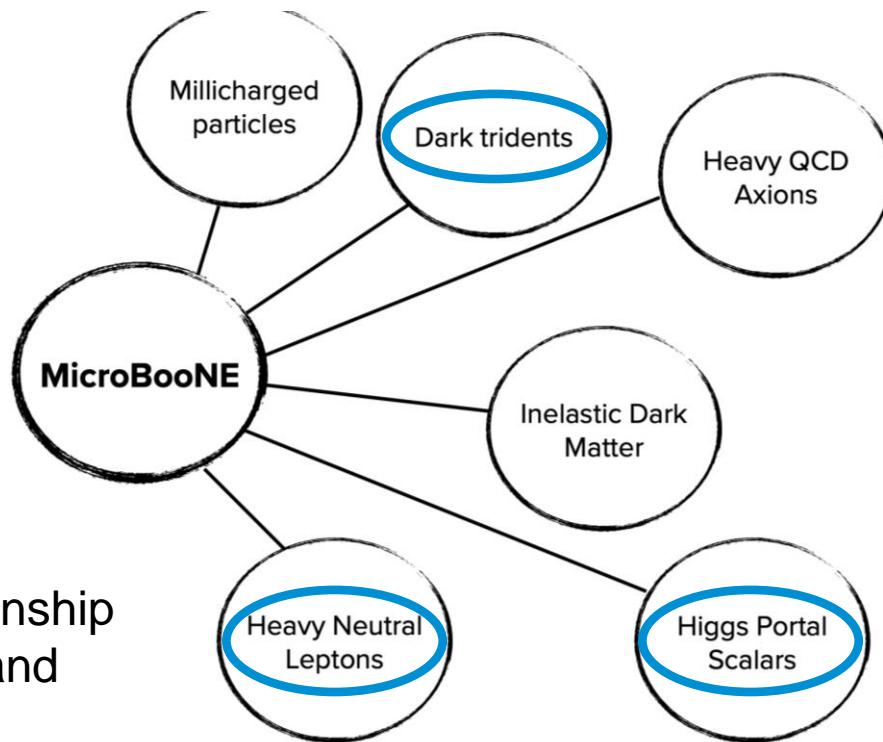
Phys. Rev. Lett. **130**, 011801 (2023)

Phys. Rev. D **106**, 092006 (2022)

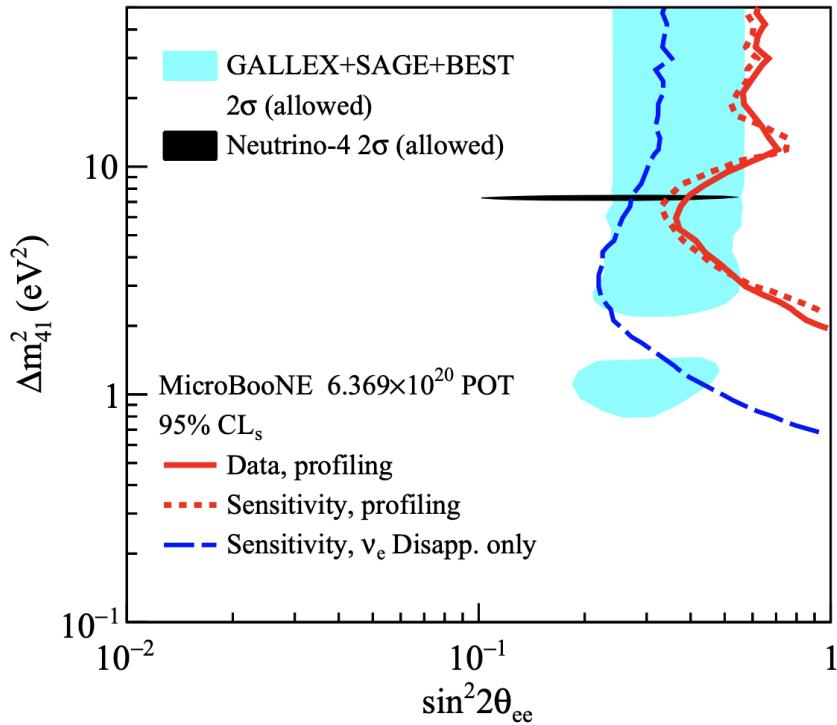
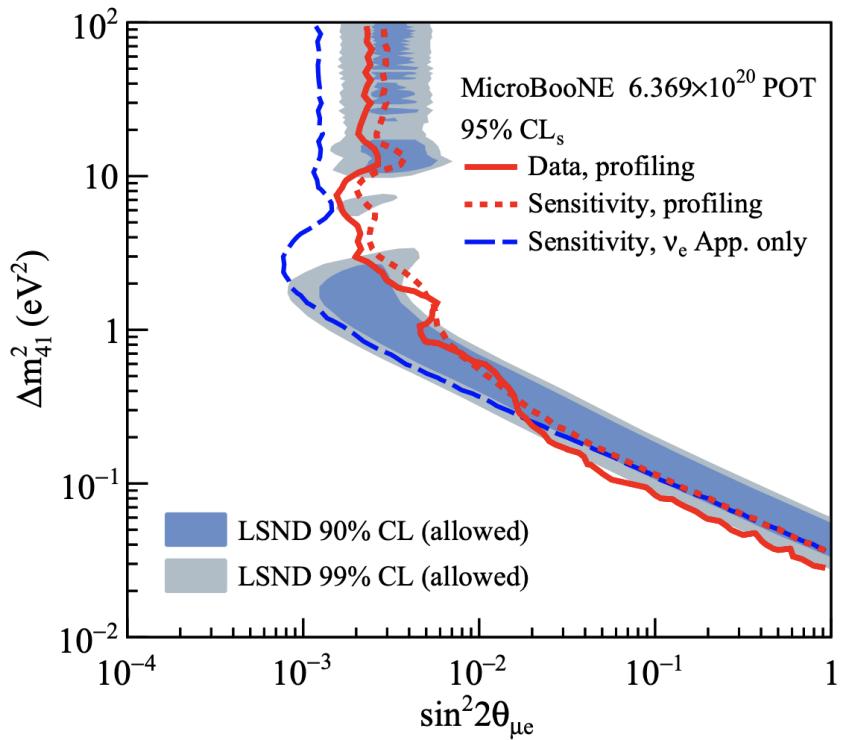
Phys. Rev. Lett. **127**, 151803 (2021)

Phys. Rev. D **101**, 052001 (2020)

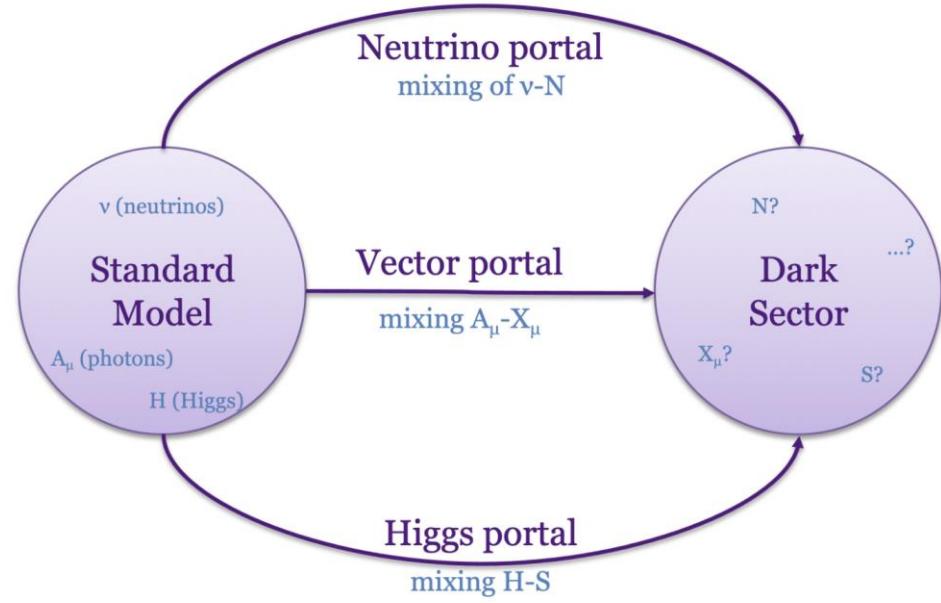
Made possible by a strong relationship
with theorists, providing models and
sensitivity estimates



Constraints on eV sterile neutrinos



Portals to dark sectors

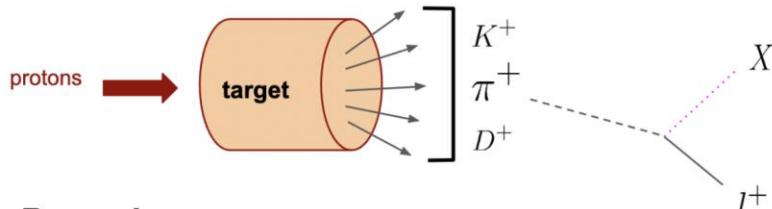


Portals to dark sectors

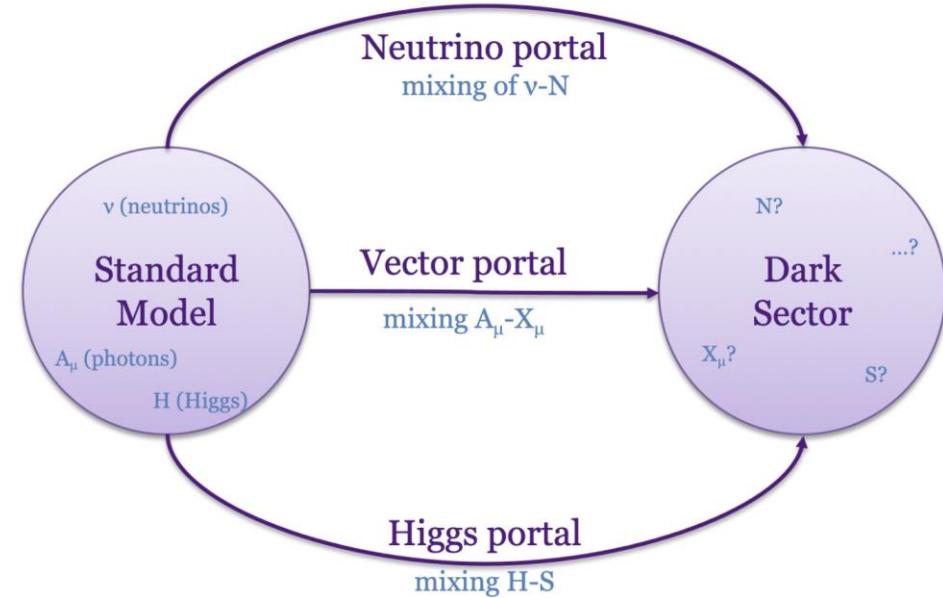
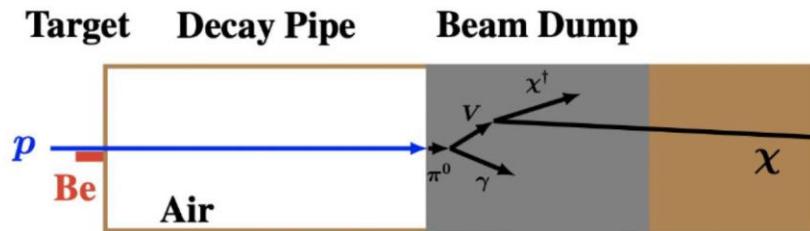
New particles can be produced from meson decays

- Large flux of charged & neutral mesons from high intensity proton beams

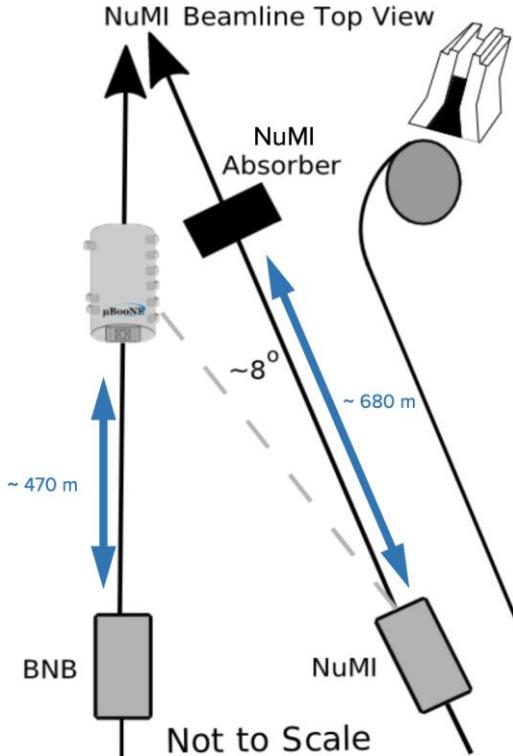
Meson decay in flight



Beam dump



Two beamlines



Booster Neutrino Beam

- 8 GeV protons
- 0.8 GeV mean neutrino energy
- On-axis

NuMI Beam

- 120 GeV protons
- 1.5 GeV mean neutrino energy
- ~8° off-axis

Heavy neutral leptons (HNLs)

Standard mixing

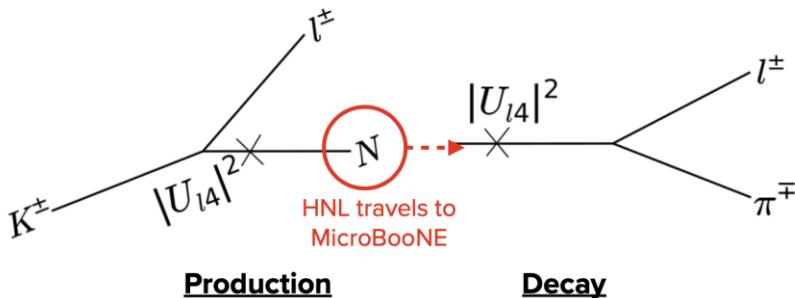
$U_{\text{PMNS}}^{\text{Extended}} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & \cdots & U_{en} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & \cdots & U_{\mu n} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & \cdots & U_{\tau n} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ U_{s_n 1} & U_{s_n 2} & U_{s_n 3} & \cdots & U_{s_n n} \end{pmatrix}$

New physics

Incorporate a new right-handed singlet that mixes with the active neutrinos

Heavy neutral leptons produced in the beam

- Decay to Standard-Model particles in the detector



Heavy neutral leptons (HNLs)

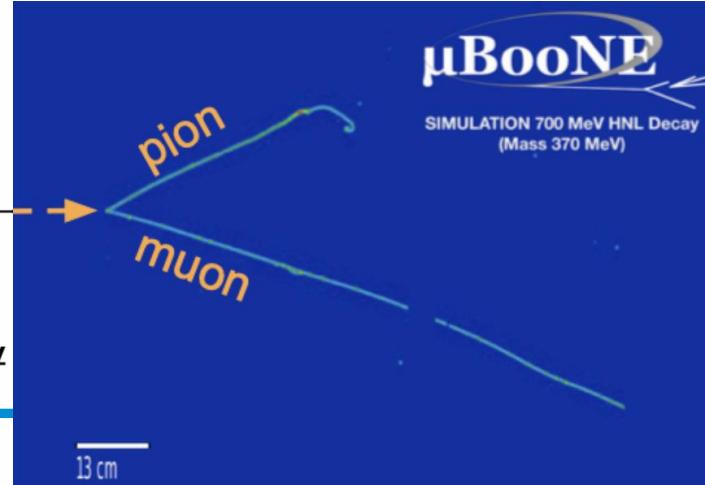
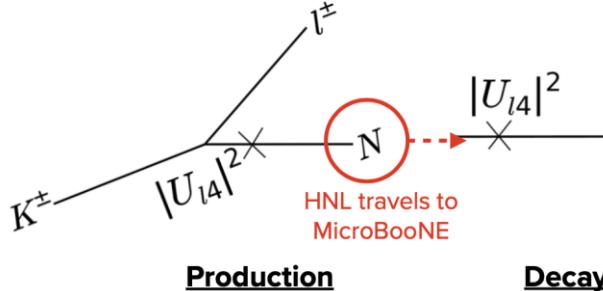
Standard mixing

$$U_{\text{PMNS}}^{\text{Extended}} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \\ \vdots & \vdots & \vdots \\ U_{s_n 1} & U_{s_n 2} & U_{s_n 3} & \cdots & U_{s_n n} \end{pmatrix}$$

$U_{\text{PMNS}}^{\text{Extended}}$ = $U_{\text{PMNS}}^{3 \times 3}$ (green box) + $U_{\text{New physics}}$ (red box)

New physics

- Production and decay rates both proportional to $|U_{l4}|^2$



Heavy neutral leptons (HNLs)

Standard mixing

$$U_{\text{PMNS}}^{\text{Extended}} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & \cdots & U_{en} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & \cdots & U_{\mu n} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & \cdots & U_{\tau n} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ U_{s_n 1} & U_{s_n 2} & U_{s_n 3} & \cdots & U_{s_n n} \end{pmatrix}$$

New physics

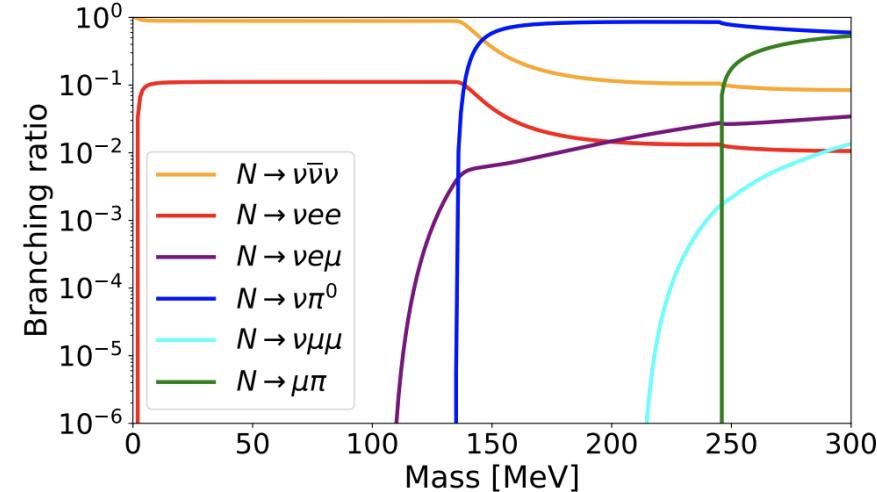
- We assume $|U_{e4}| = |U_{\tau 4}| = 0$ and set limits on $|U_{\mu 4}|$
- We focus on production via $K \rightarrow \mu N$

Heavy neutral leptons (HNLs)

Standard mixing

$$U_{\text{PMNS}}^{\text{Extended}} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & \cdots & U_{en} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & \cdots & U_{\mu n} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & \cdots & U_{\tau n} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ U_{s_n 1} & U_{s_n 2} & U_{s_n 3} & \cdots & U_{s_n n} \end{pmatrix}$$

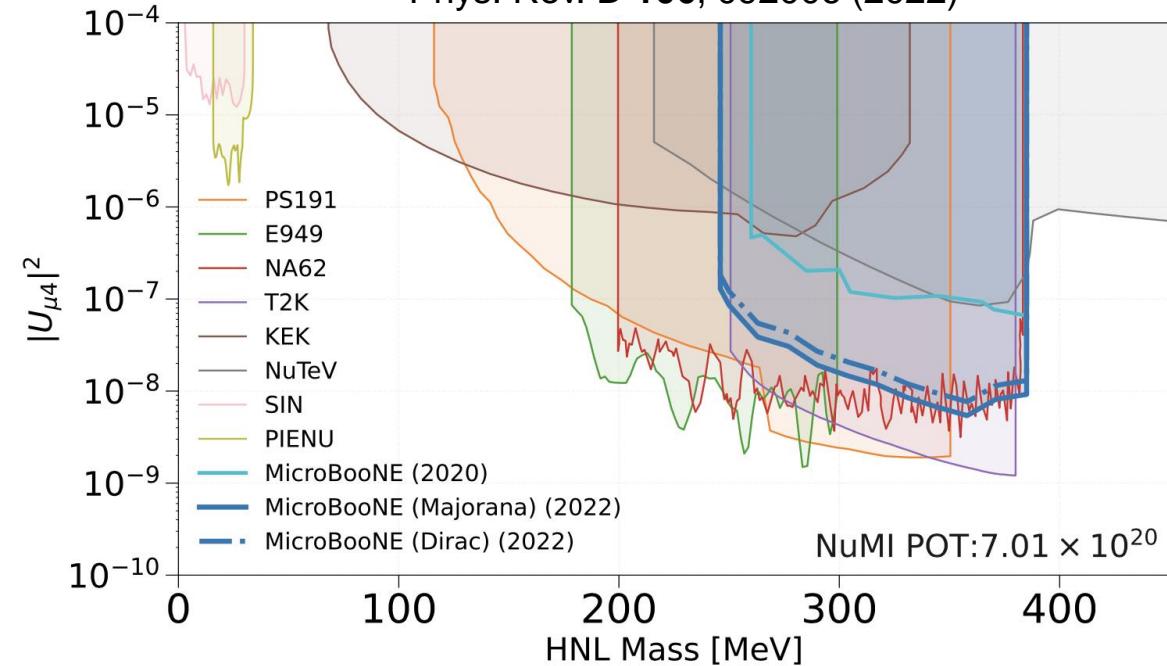
New physics



- We assume $|U_{e4}| = |U_{\tau 4}| = 0$ and set limits on $|U_{\mu 4}|$
- We focus on production via $K \rightarrow \mu N$
- Most likely decay channel depends on m_N

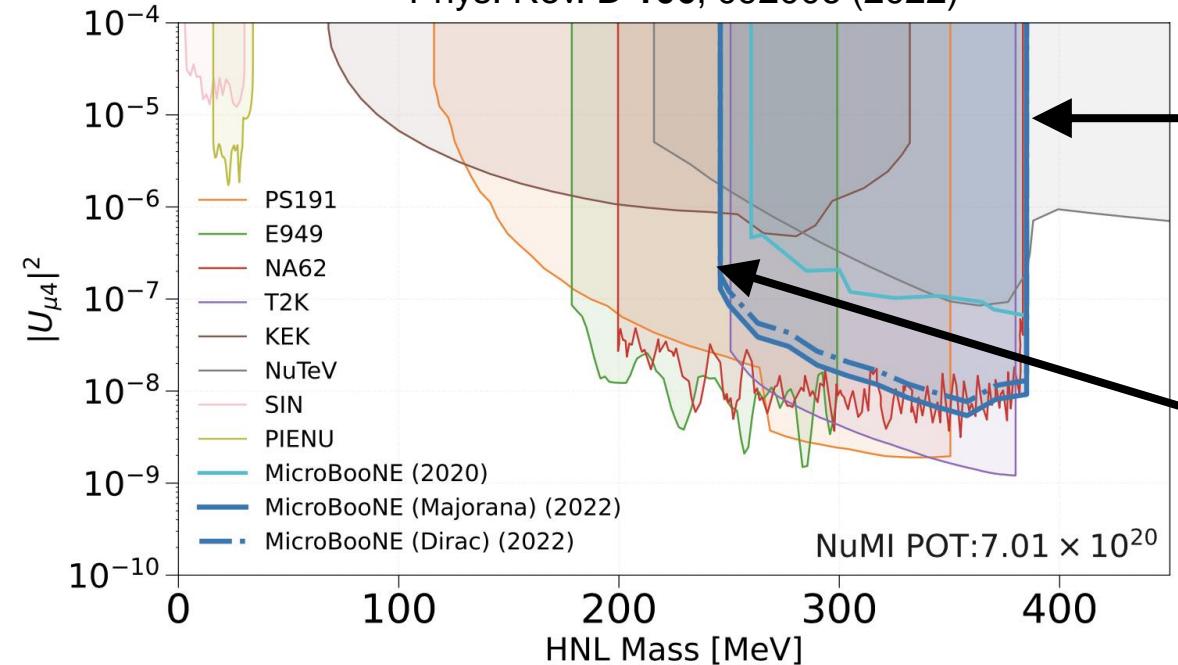
Search for $N \rightarrow \mu^\pm \pi^\mp$

Phys. Rev. D **106**, 092006 (2022)



Search for $N \rightarrow \mu^\pm \pi^\mp$

Phys. Rev. D **106**, 092006 (2022)

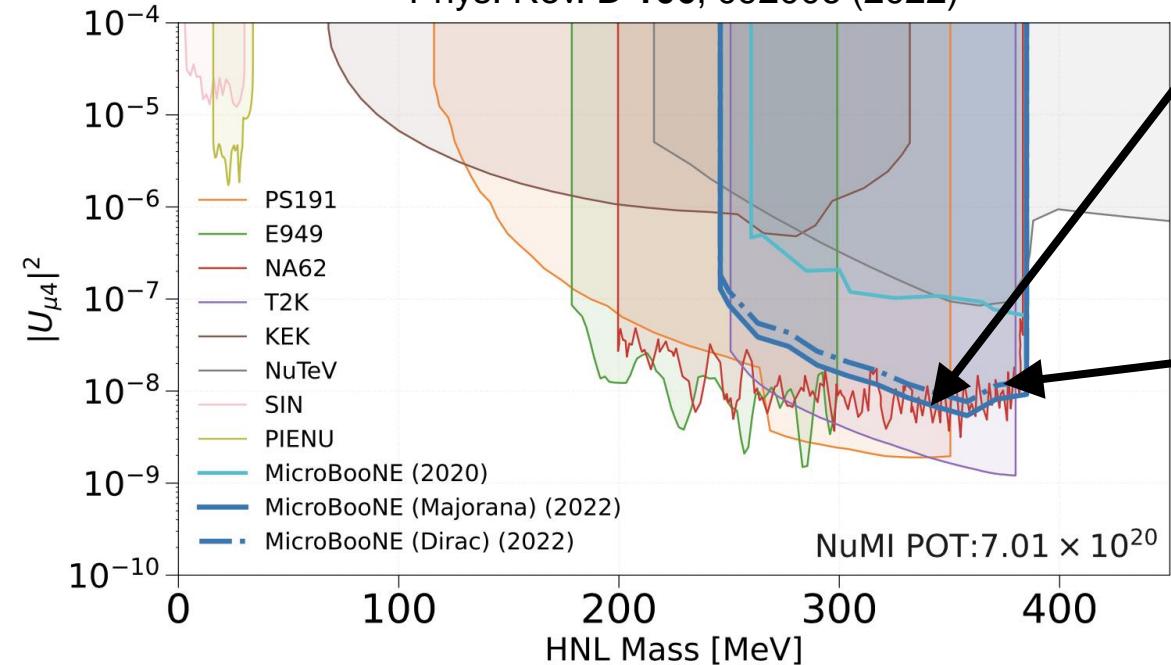


Kinematic limit for $K \rightarrow \mu N$ production

Kinematic limit for $N \rightarrow \mu \pi$ decay

Search for $N \rightarrow \mu^\pm \pi^\mp$

Phys. Rev. D **106**, 092006 (2022)



Majorana HNL

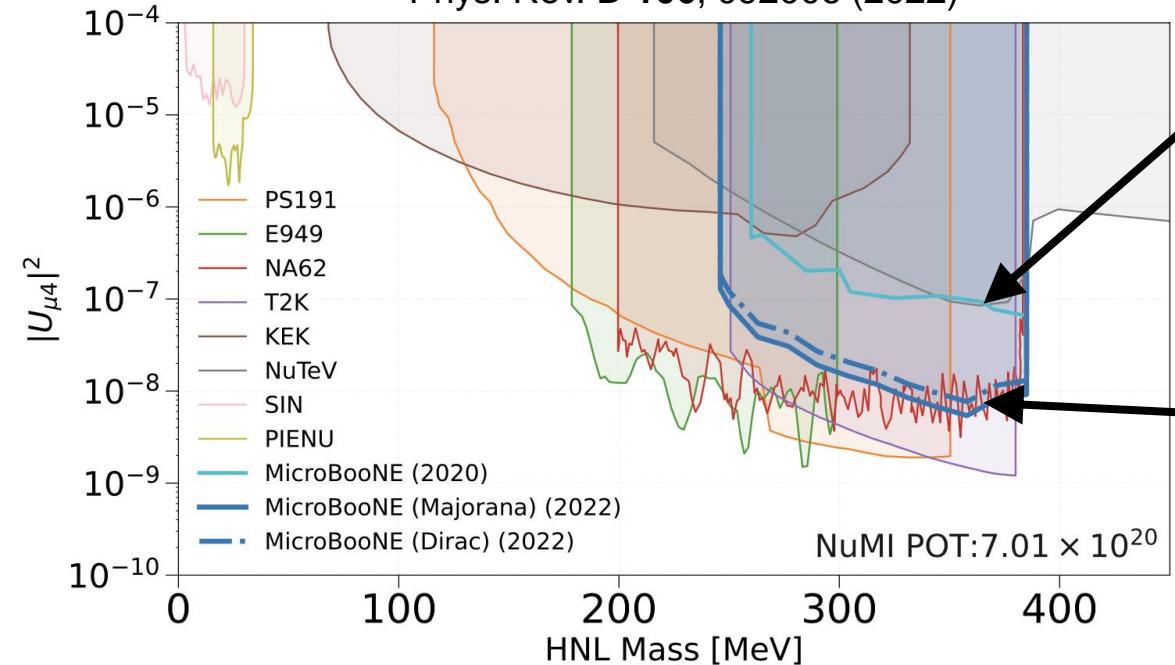
- $N \rightarrow \mu^+ \pi^-$ and $N \rightarrow \mu^- \pi^+$
- Isotropic decays (summed over both channels)

Dirac HNL

- $N \rightarrow \mu^- \pi^+$ only: factor of two lower rate
- Non-isotropic decays (negligible effect)

Search for $N \rightarrow \mu^\pm \pi^\mp$

Phys. Rev. D **106**, 092006 (2022)



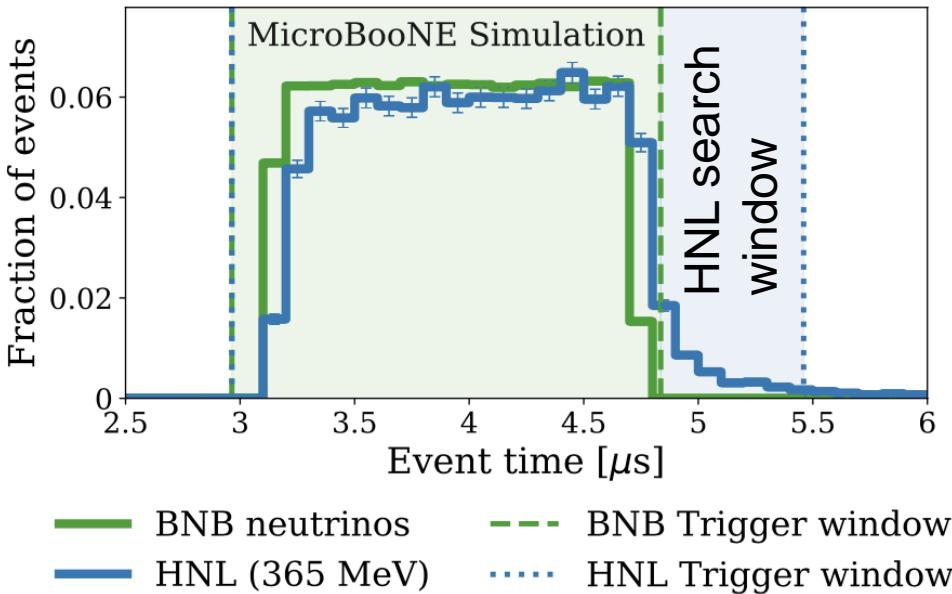
BNB search

Phys. Rev. D **101**, 052001 (2020)

NuMI search

Phys. Rev. D **106**, 092006 (2022)

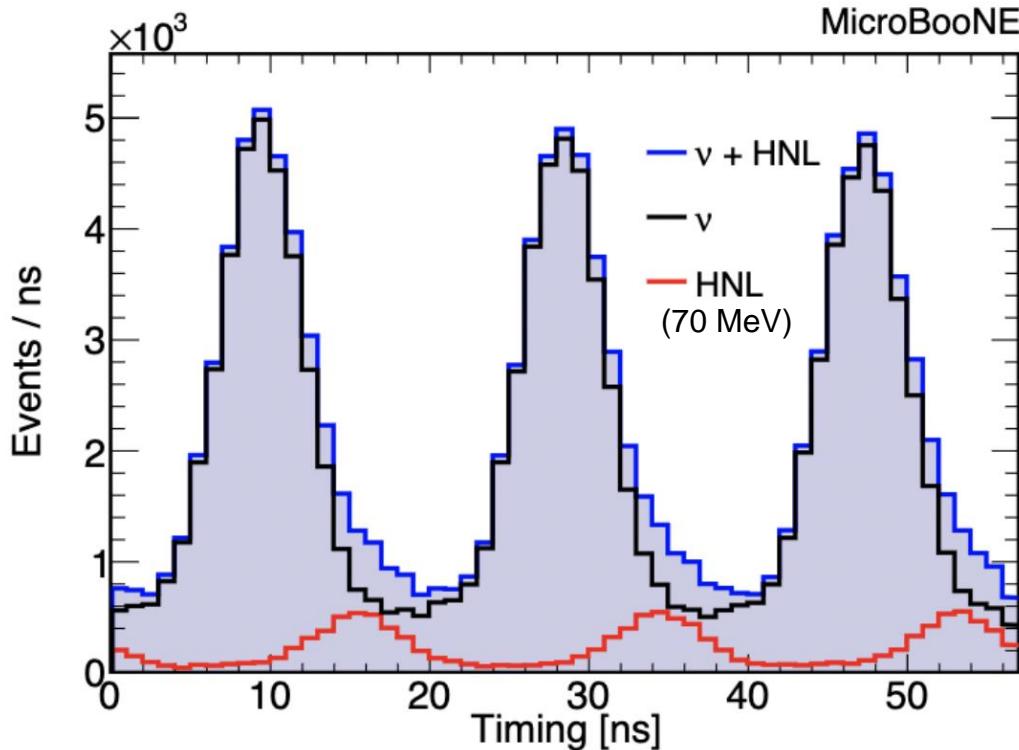
BNB search: late-arrival trigger



Used the fact that heavy HNLs arrive later than light neutrinos

- Implemented a dedicated late-arrival trigger to eliminate regular neutrino interactions
- Uses a 0.6 μs window after the usual BNB trigger

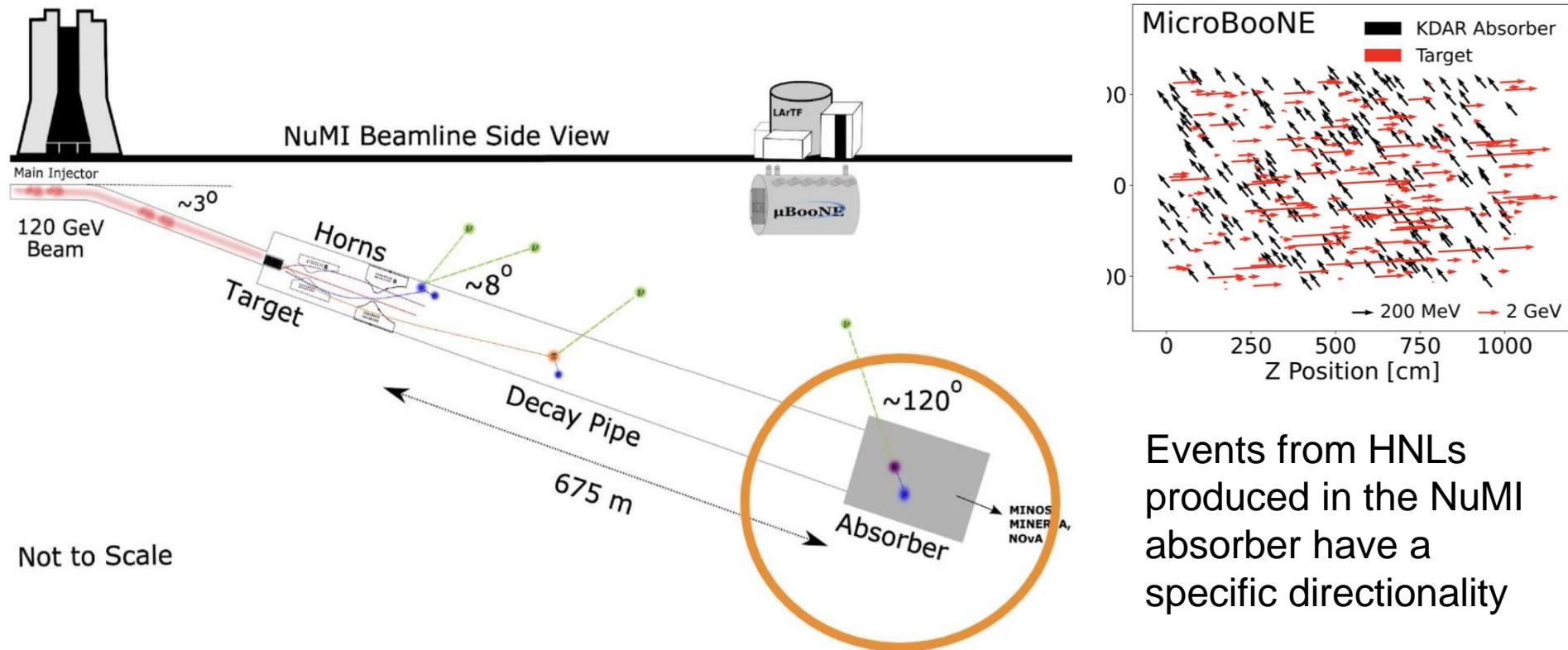
Potential upgrade with ns timing



Phys. Rev. D **108**, 052010 (2023)

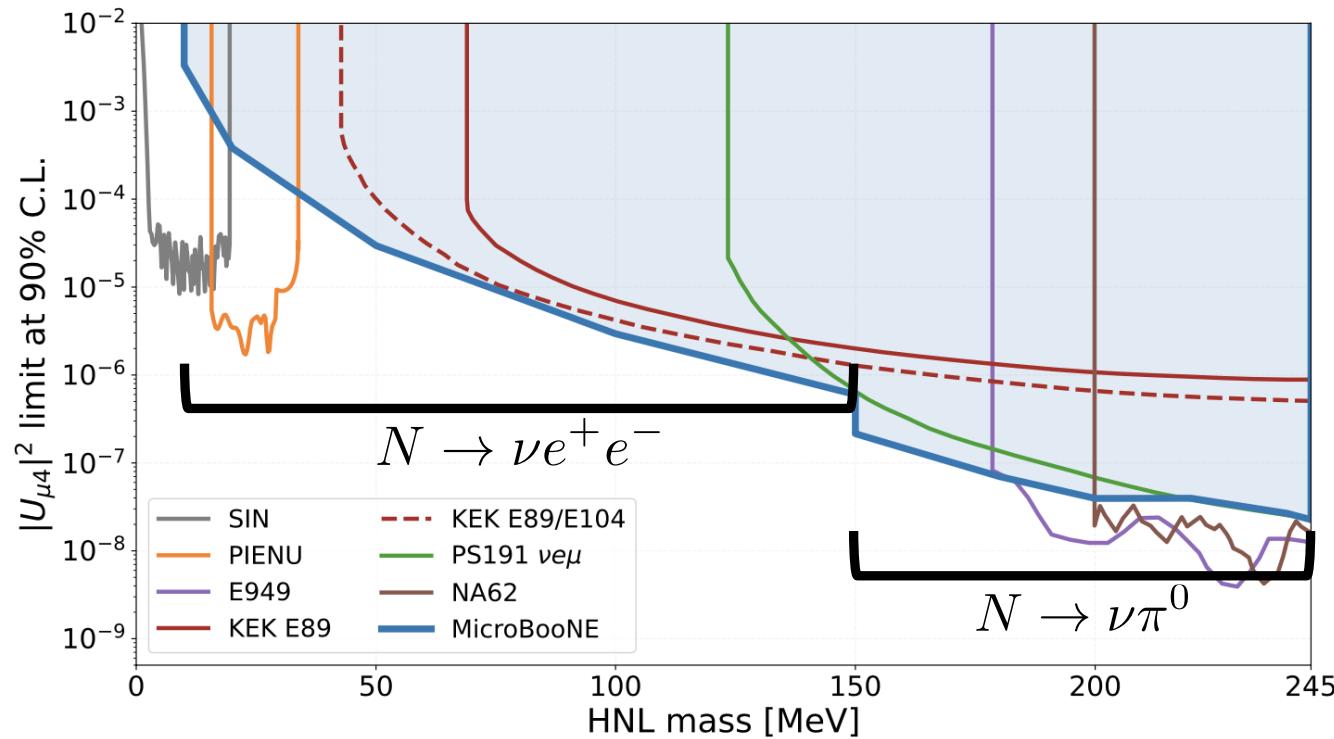
Dante Totani's talk

NuMI search: off-axis

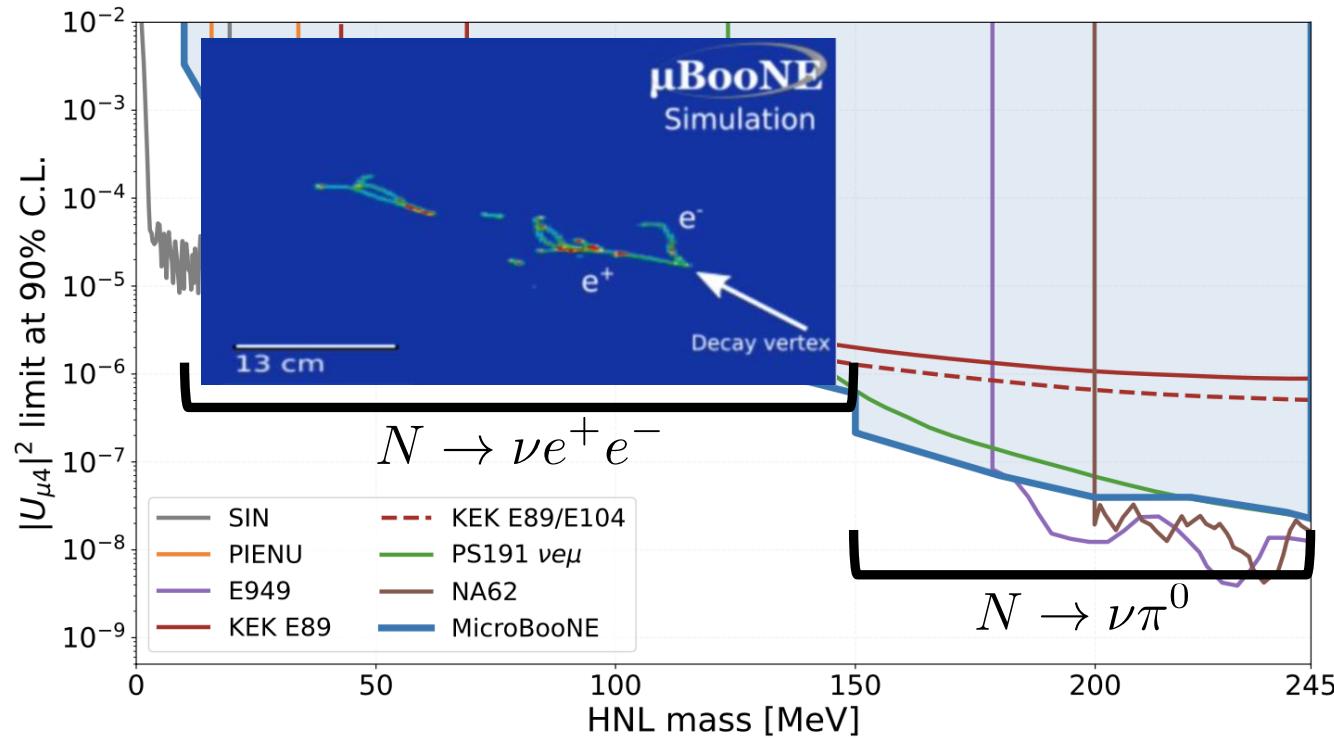


Events from HNLs produced in the NuMI absorber have a specific directionality

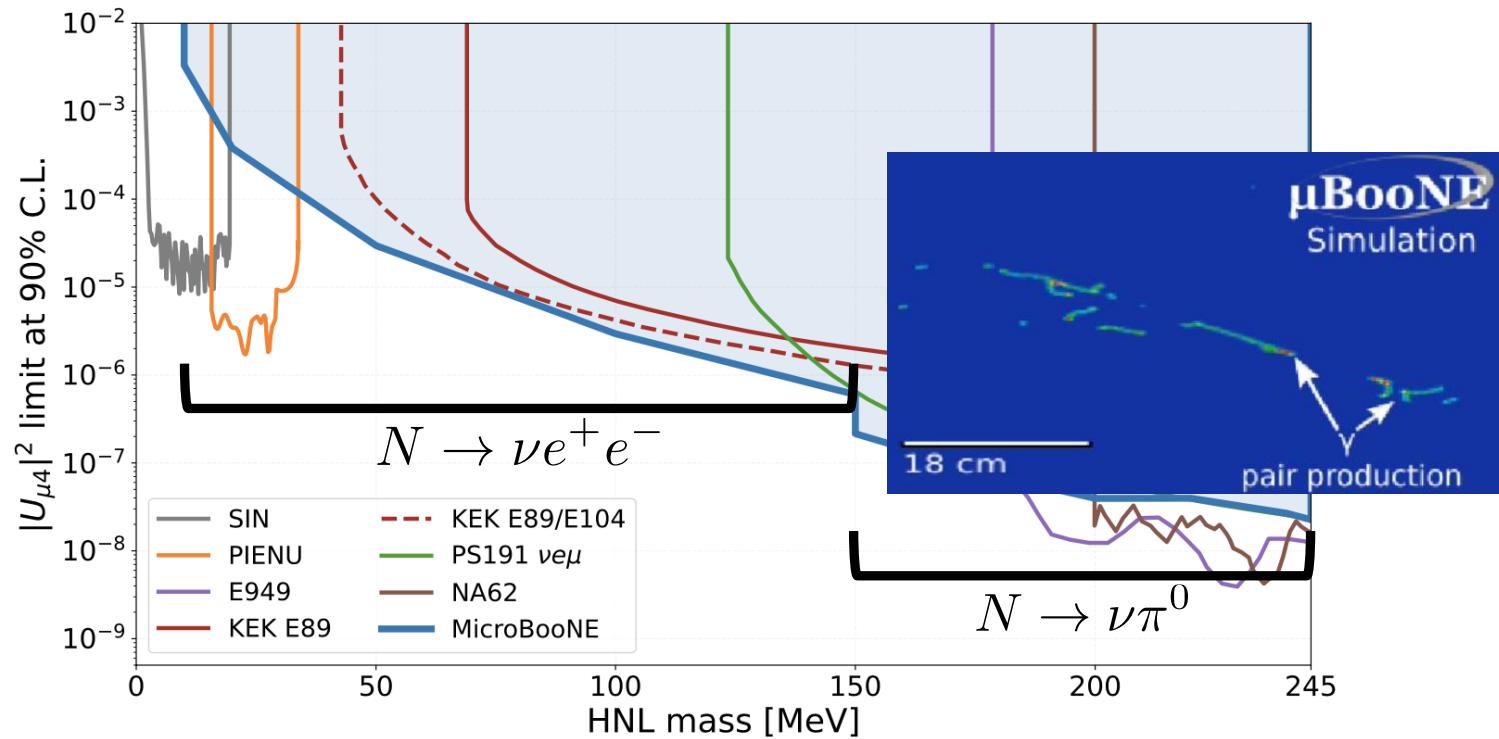
Lighter heavy neutral leptons



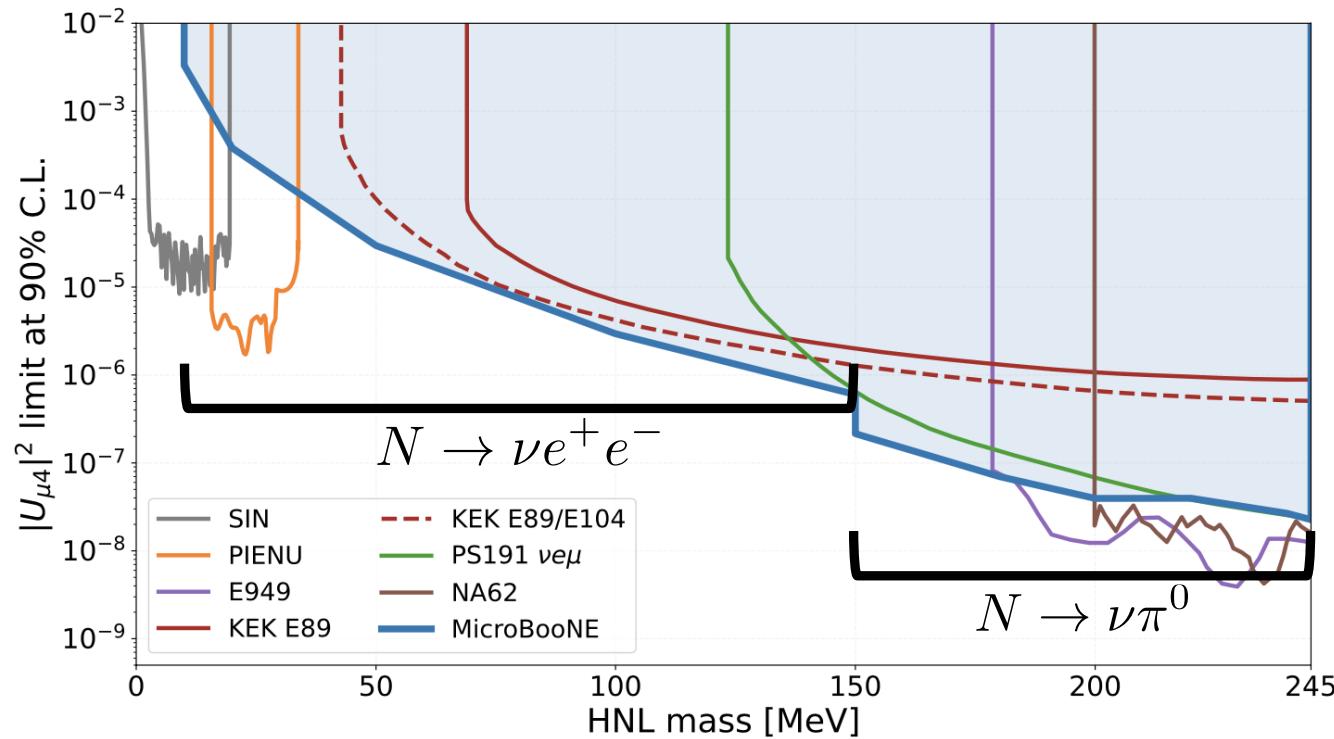
Lighter heavy neutral leptons



Lighter heavy neutral leptons

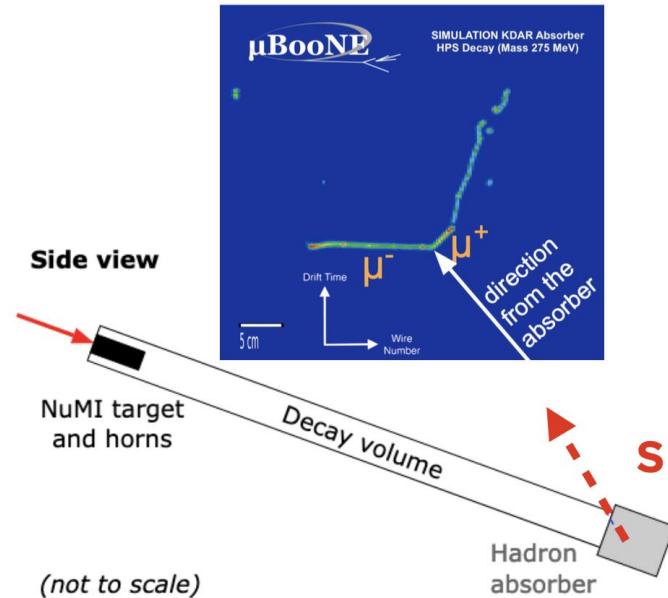
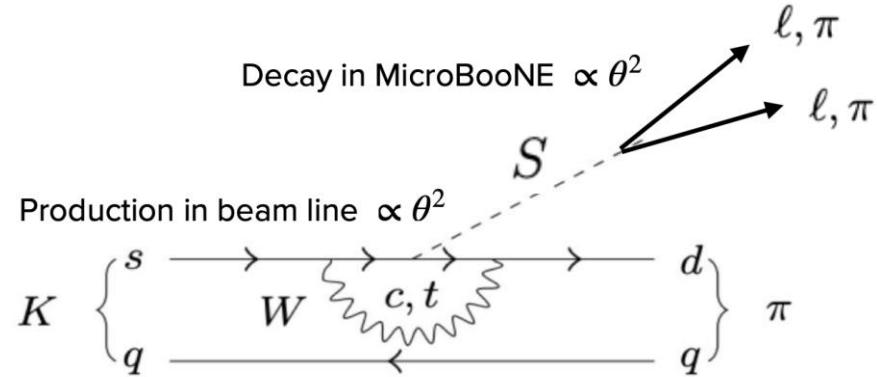


Lighter heavy neutral leptons

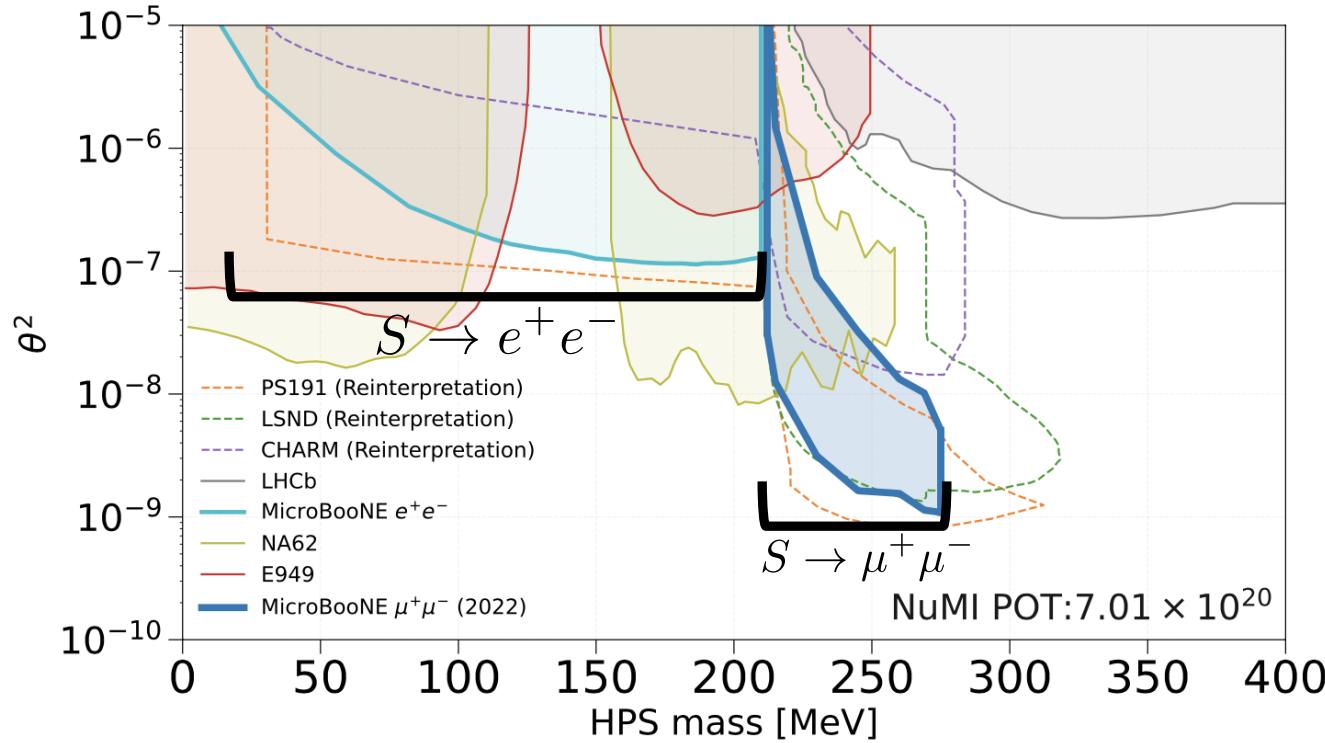


Higgs portal scalars

- Neutral singlet scalar, S
- Mixes with the Higgs, parameterized by an angle θ
- Produced in meson decays in the beam
- Decays into Standard-Model particles in the detector



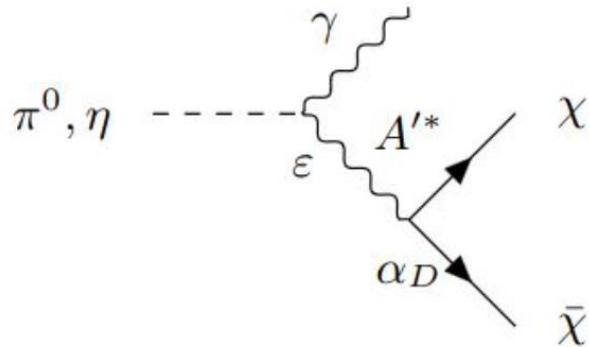
Higgs portal scalars



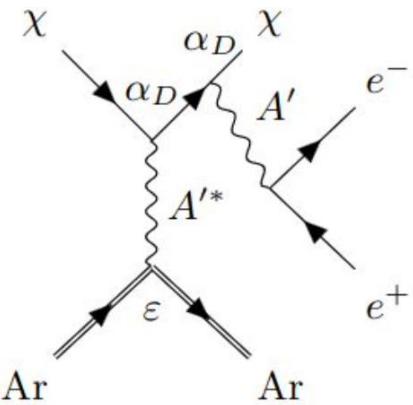
Dark tridents

arXiv:2312.13945

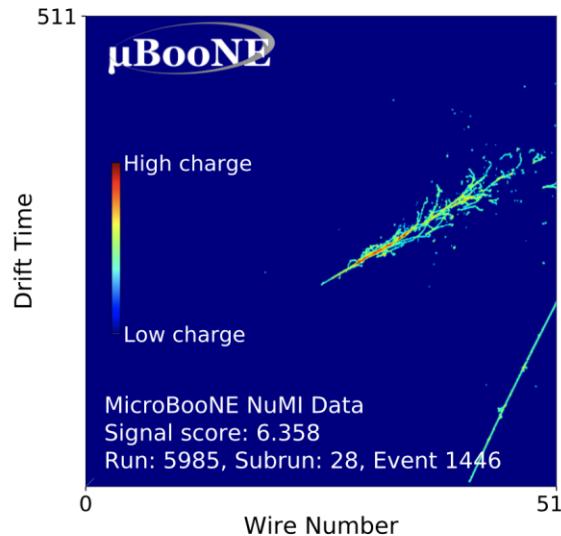
Production in
the NuMI beam



Interaction in
the detector



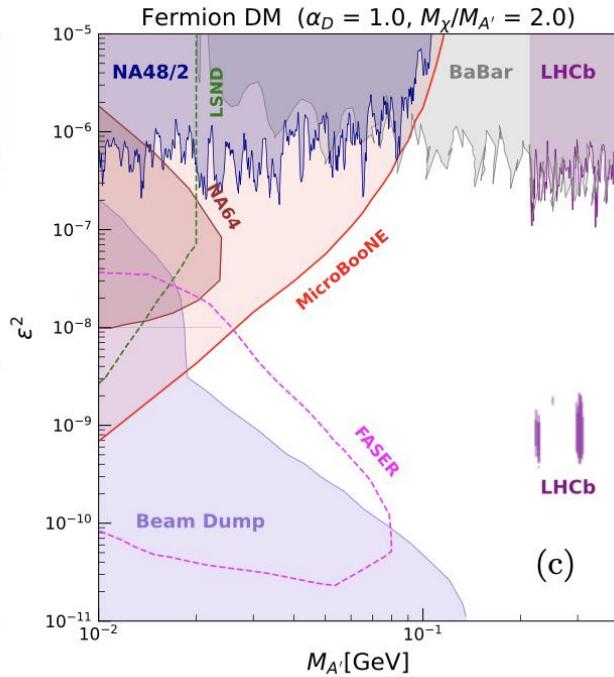
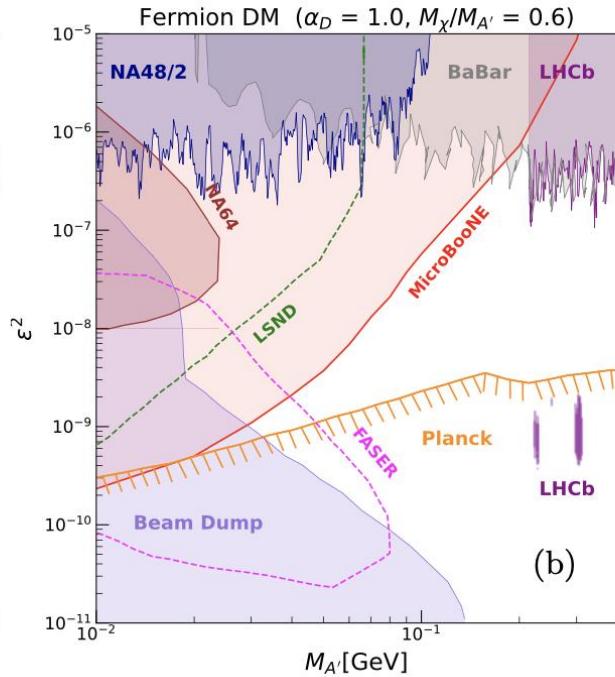
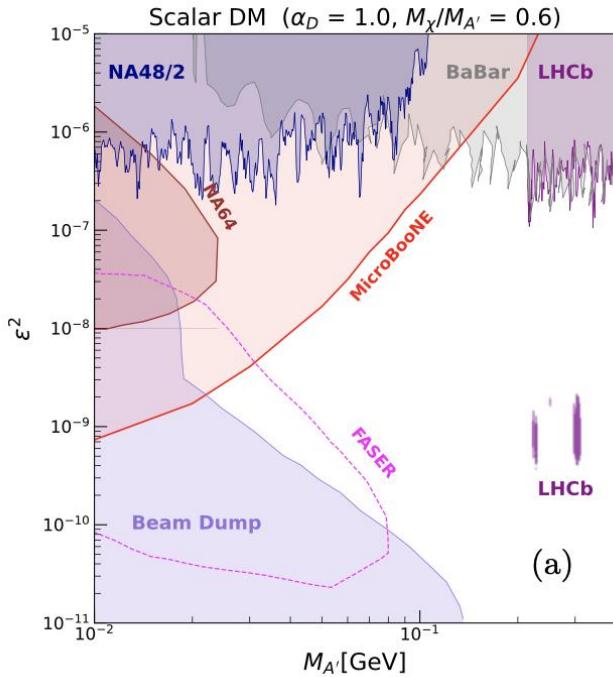
- Dark matter particles, χ , produced in decays of neutral mesons
- χ particles interact in the detector, mediated by a dark photon
- Free parameters are ε , α_D and M_χ/M_A



Uses pixel-based
deep learning
techniques

Dark tridents

arXiv:2312.13945



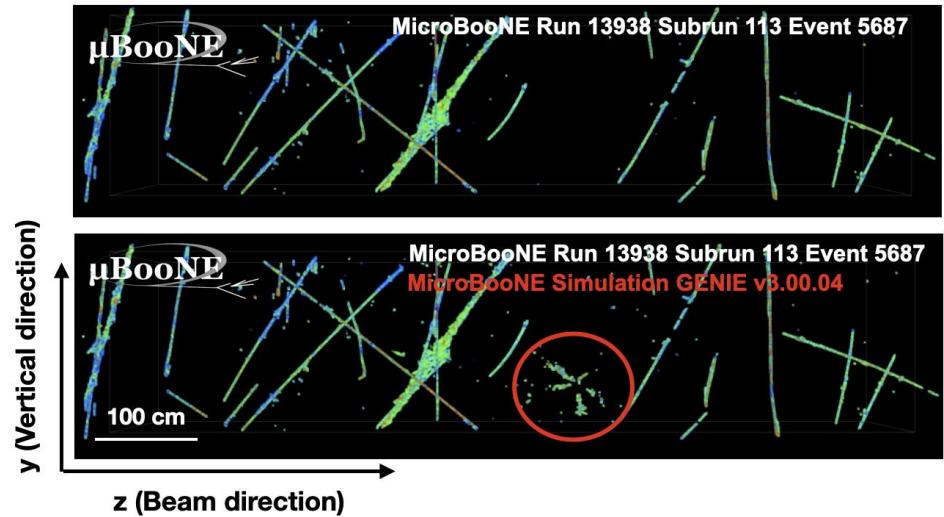
Neutron-antineutron oscillations

Neutron transforms into an antineutron

- Annihilates with a neighbouring neutron
- Produces a star-like topology of pions

We use pixel-based deep learning techniques to identify the topology

- Achieve 70% signal efficiency
- Our techniques can improve DUNE's published efficiency for neutron-antineutron searches by a factor of 7



Summary

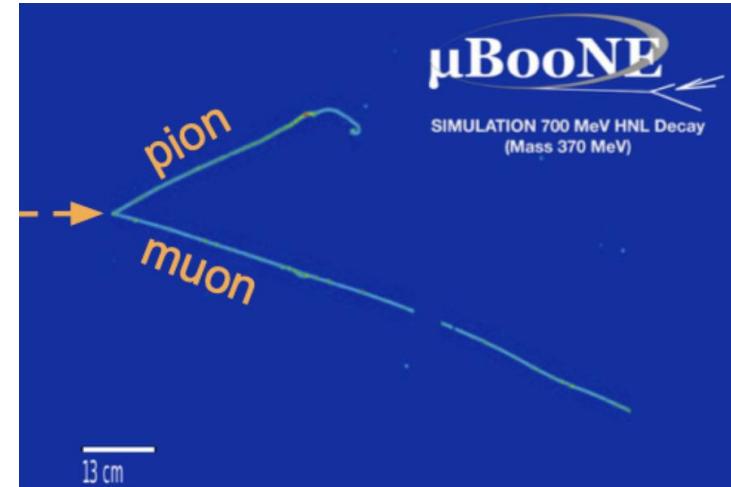
Searches for BSM particles have been an integral part of MicroBooNE's physics programme

- Heavy neutral leptons
- Higgs portal scalars
- Dark tridents
- Light sterile neutrinos

More analyses are in the pipeline

- e^+e^- production, axion-like particles, millicharged particles

The full SBN programme, and DUNE, will be able to expand on these searches



We're eager for more ideas of models we can test with our data!